

Imaging the Pleiades with VEGA / CHARA

March 15th, 2016 F. Millour, A. Meilland, A. Domiciano de Souza, P. Stee



















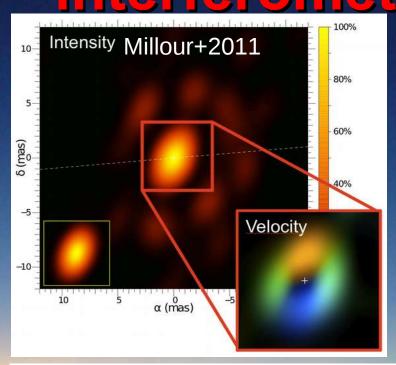


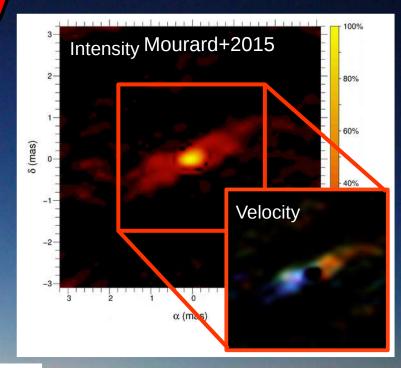




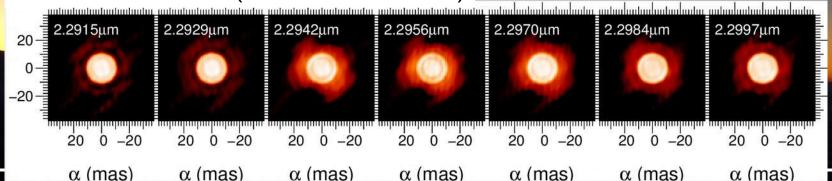


Image cubes with interferometry





CO bandhead on Mira (Fabas et al. submitted)





Why observing the Pleiades?

- Open cluster with all stars born at the same time
 - Evaluate the kinematics of the primordial molecular cloud by measuring rotation velocities
 - Parent nebula topology by measuring spin-axes orientation of several stars
- Image the disk kinematics of several active Be stars
 - 4 stars are in active Be phase: Alcyone, Pleione, Electra, and Merope















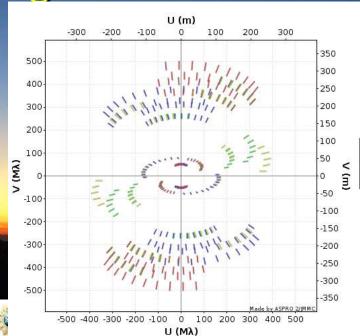


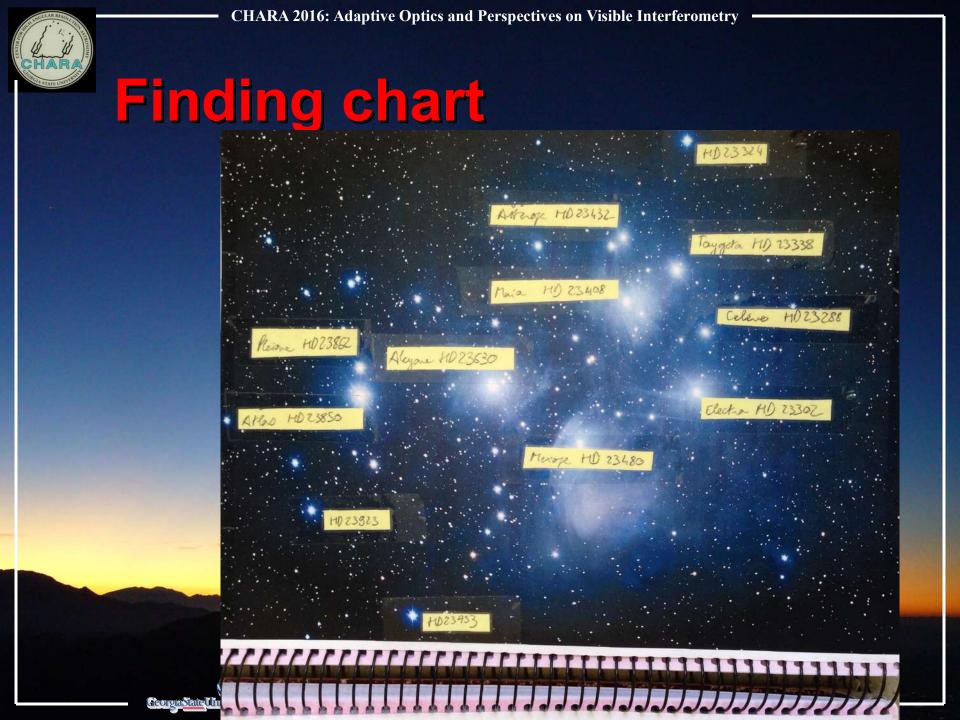


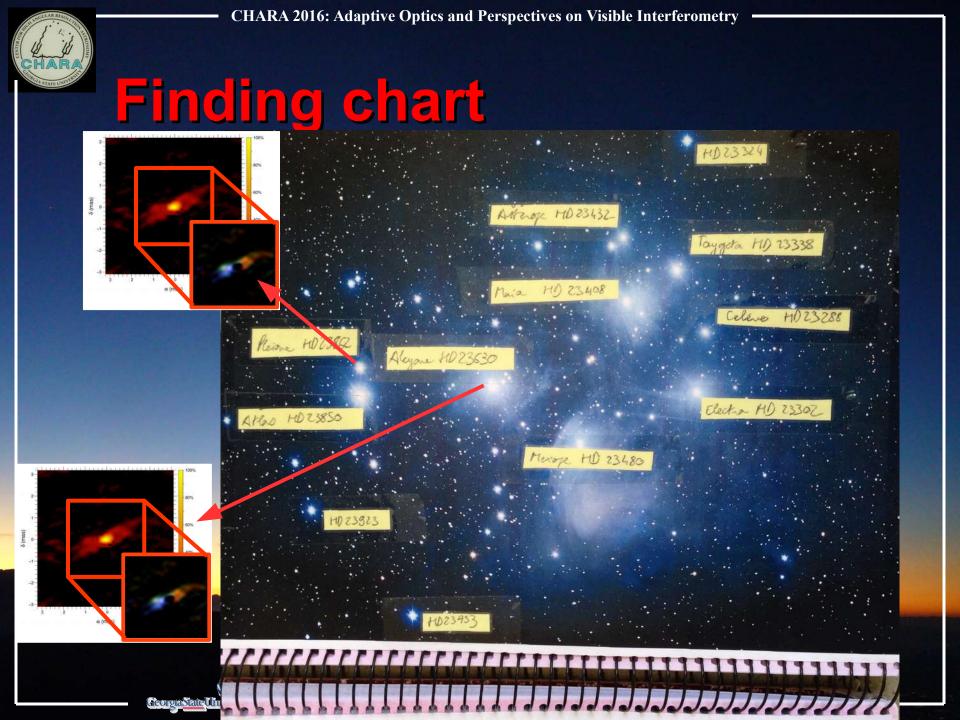
Imaging potential

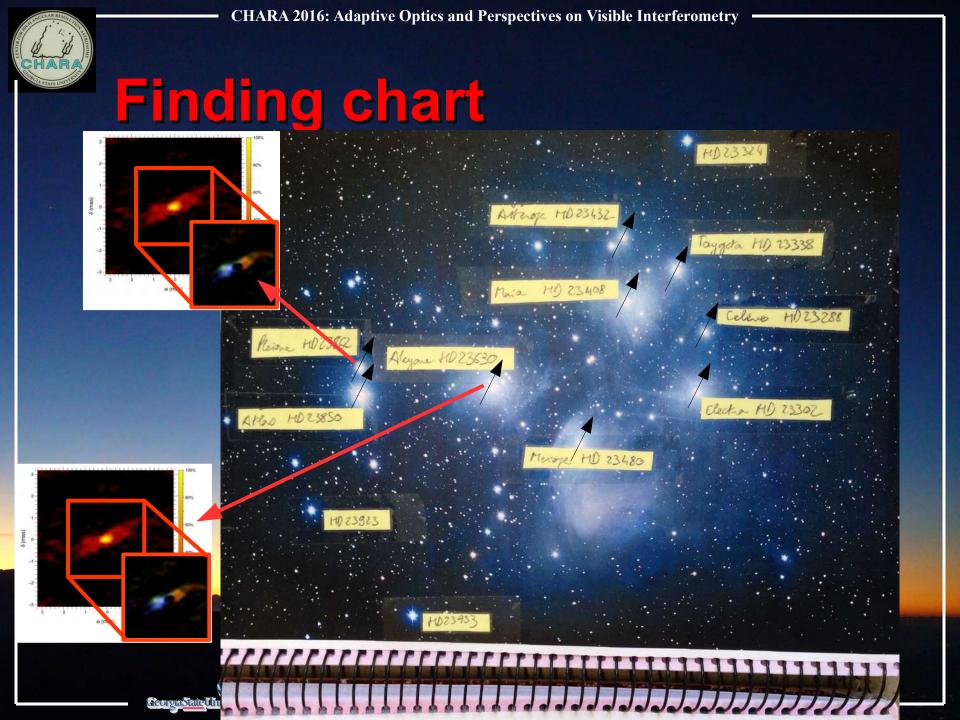
 Alcyone & Pleione present strong-enough emission lines and large-enough discs to repeat the imaging experiment of Phi Per

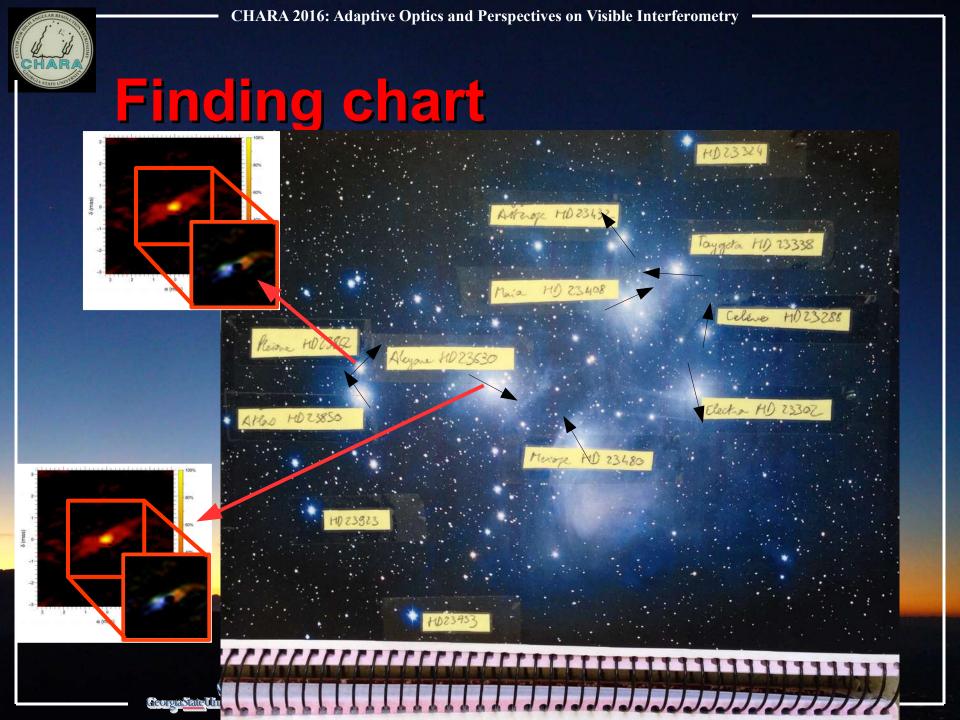
• (u,v) coverage is OK











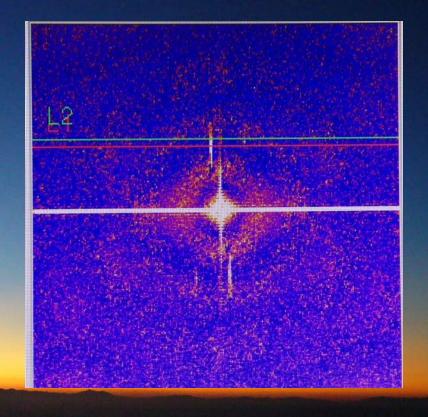


Observations done

4T at the beginning

TF / AC Display-264 EraseFT1 EraseFT2 EraseFT3 DispFT1 DispFT2 DispFT3

3T at the end









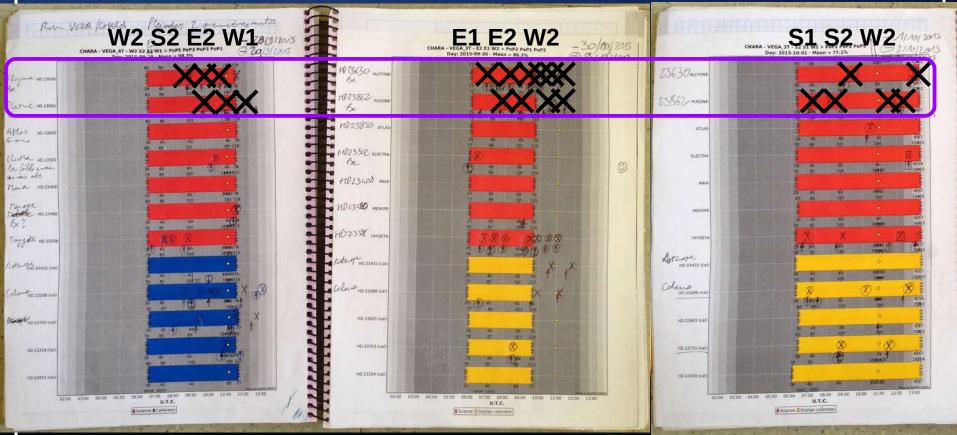






Observations done

28/09/2015 29/09/2015 30/09/2015 01/10/2015 02/10/2015 09/10/2015



















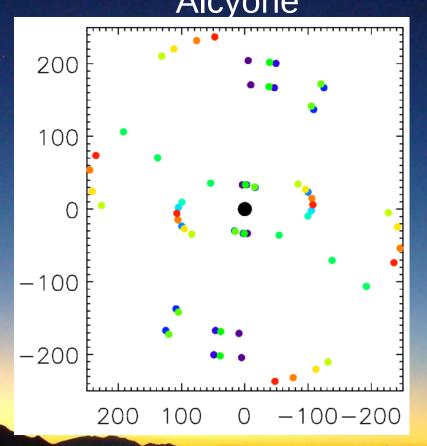




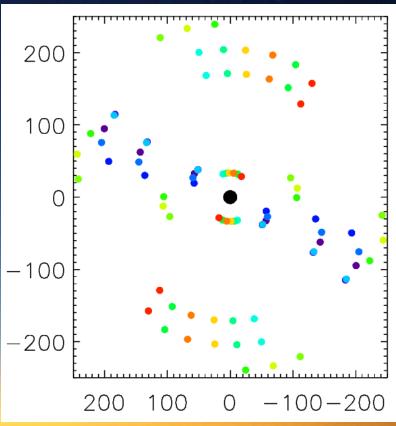


(u,v) coverage Alcyone





Pleione

















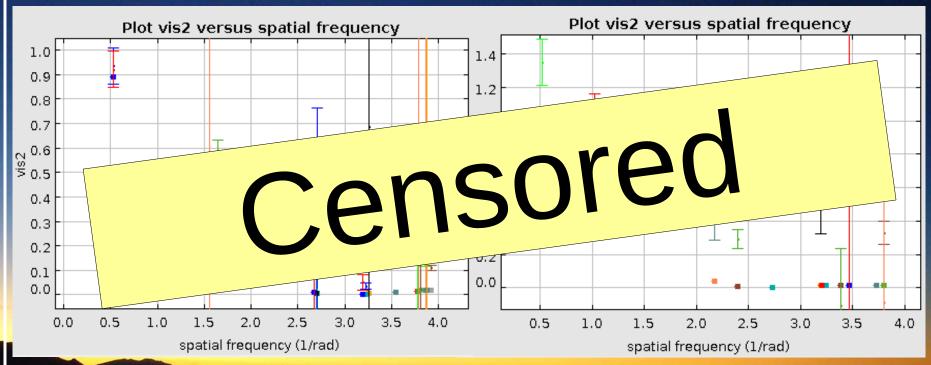






(preliminary) visibilities





More work to be done to calibrate properly the visibiltiies LitPro fit: bad chi2













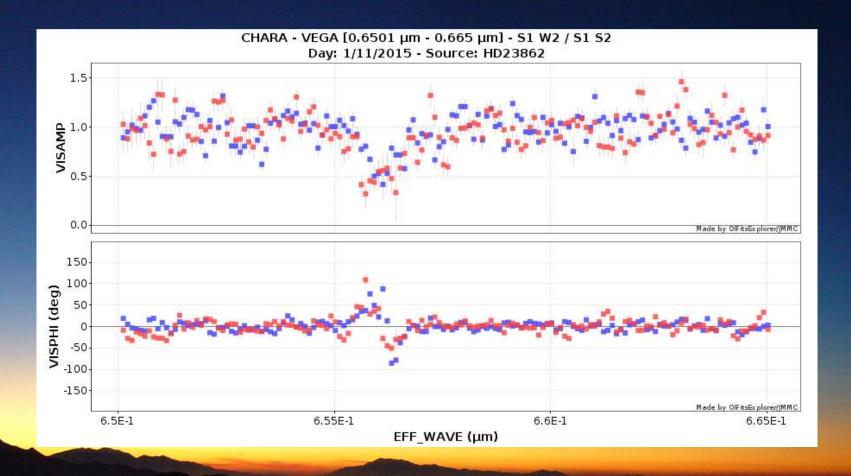








Pleione















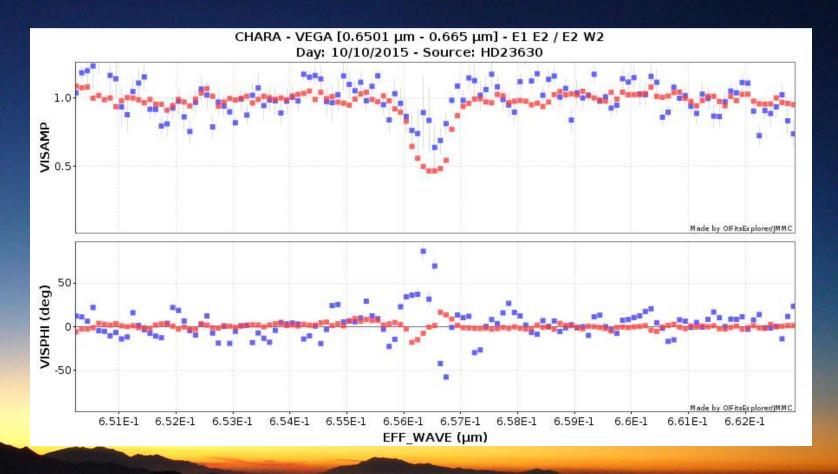








Alcyone























Thanks!

- Done: observing
- Done: data reduction
- TBD: reduce the data
- TBD: data not reduced good enough, repeat
- TBD: image reconstruction
- TBD: model fitting
 - TBD: publish

















